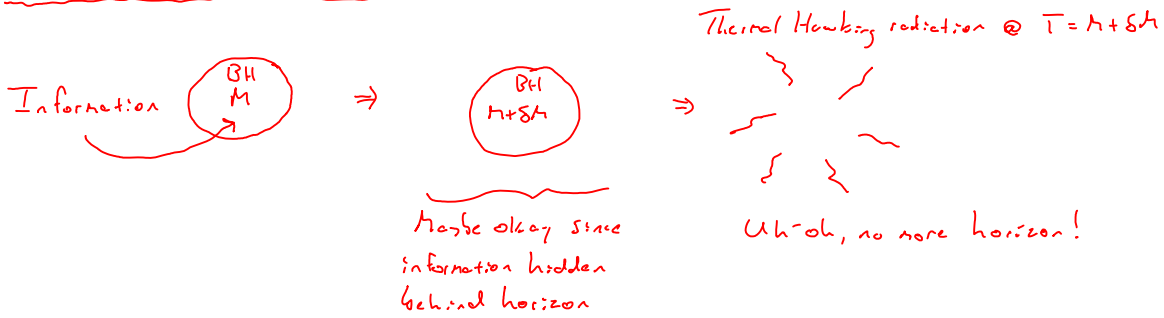


Brief review of the BIP



If information lost: Non-unitary evolution which means quantum anything gets broken.
If Hawking radiation is wrong: Something off about effective field theory, equivalence principle, etc.

Partial resolution via AdS/CFT ⇒ Unitarity must be preserved!
almost anything gravity non-gravity QFT

But how? While this, and his own similar argument, led Hawking to concede the bet... the details were still unclear.

So folks kept thinking.

Since these ideas hover around entropy and information, let's sharpen these notions a bit.

Heuristically: Entropy is a measure of what we can know, but choose not to know about a system. Entropy is a physically meaningful (and impactful) quantity because the process of gaining knowledge is achieved through physical investigation, i.e. detection and we know that this has physical consequences.

Classical Thermodynamic Entropy: For macroscopic systems $dE = T dS + \Omega dJ + \Phi dQ$
 (under reversible changes from eq to eq) Δ excess Δ entropy Δ in ang mom Δ in elec charge
 \int temp \int mag. pot. \int elec. pot.
heat transferred

$E, T, S, \Omega, J, \Phi, Q$ are macro. state variables

Statistical Entropy: Given a coarse-grained description (in terms of average quantities only for example) there are usually many fine-grained resolutions compatible.

If the i th compatible resolution (microstate) has probability p_i to be the true state of the system then $S_S \propto -\sum_i p_i \ln p_i$. Note for $p_i = 1, p_{i \neq i} = 0$
 $S_S = 0$

The language we use is that for a coarse-grained description, there is an ensemble of resolutions such that $\langle \text{What we know} \rangle = \sum_i p_i |i \text{th resolution} \rangle$

Microcanonical: Fix $E + N \Rightarrow$ equal prob. $\Rightarrow p_i = \frac{1}{\#} \Rightarrow S_S = k_B \sum_i \frac{1}{\#} \ln \# = k_B \ln \#$
 For $\# = 1 \Rightarrow S_S = 0$

Canonical: Fix $T + N \Rightarrow p_i = \frac{1}{Z} e^{-E_i/k_B T}$ w/ $Z = \sum_i e^{-E_i/k_B T}$ so $\sum_i p_i = 1$

$S_S = k_B (\ln Z + \frac{\langle E \rangle}{T})$ for $T \rightarrow 0$ (only E_0 contributes) $\Rightarrow S_S \rightarrow 0$

In all cases S is extensive, i.e. $S(A \cup B) = S(A) + S(B)$ (volume scaling)

Von Neumann Entropy: In a quantum system there is a limit to how much we can possibly know, i.e. uncertainty, e.g. $[\hat{x}, \hat{p}_x] \neq 0$. This means in particular that even if we measure (know) everything allowed we can still be in a superposition $|\psi\rangle = \sum_i c_i |E_i\rangle$, but this is the state of the system so should have $S=0$. $\sum_i |c_i|^2 = 1$

To accommodate this define $\rho = |\psi\rangle\langle\psi|$ then $S_{vN} = -\text{Tr}(\rho \ln \rho)$

E.g. $|\psi\rangle = a |\uparrow\rangle_A |\downarrow\rangle_B + b |\downarrow\rangle_A |\uparrow\rangle_B$ w/ $|a|^2 + |b|^2 = 1$
 $\rho = |\psi\rangle\langle\psi| = (a |\uparrow\rangle_A |\downarrow\rangle_B + b |\downarrow\rangle_A |\uparrow\rangle_B)(a^* \langle\uparrow|_A \langle\downarrow|_B + b^* \langle\downarrow|_A \langle\uparrow|_B)$
 $= \begin{pmatrix} aa^* & ab^* \\ ba^* & bb^* \end{pmatrix} \Rightarrow \rho^2 = \rho$ (always for pure state)

$S_{vN} = -\text{Tr}(\rho \ln \rho) = -\text{Tr}(\rho \ln \rho^2) = -\text{Tr}(2\rho \ln \rho) = -2 \text{Tr}(\rho \ln \rho) \Rightarrow 1=2$ unless $S_{vN} = 0$

If on the other hand we do not measure everything we can, then the evolution of our coarse graining can be represented by a collection of (possibly superposition) states $|\psi_j\rangle$ and we then have what is equivalent to an ensemble but in quantum parlance called a "mixed state". We cannot describe it w/ a state vector, but the density matrix comes in handy here: $\rho = \sum_j p_j |\psi_j\rangle\langle\psi_j|$

Where things can get interesting is:

Now suppose we choose to know less by only knowing about subsystem A. To get to this description from $|\psi\rangle$ above, we trace over the state space of B. This leads to the reduced density matrix:

$\rho_A = \text{Tr}_B \rho = \sum_B \langle\downarrow|_B \rho |\downarrow\rangle_B + \sum_B \langle\uparrow|_B \rho |\uparrow\rangle_B = |b|^2 |\downarrow\rangle_A \langle\downarrow|_A + |a|^2 |\uparrow\rangle_A \langle\uparrow|_A$
 $= \begin{pmatrix} |a|^2 & 0 \\ 0 & |b|^2 \end{pmatrix} \Rightarrow \rho_A^2 \neq \rho_A$

$S_{vN,A} = -\text{Tr}(\rho_A \ln \rho_A) = -\text{Tr}(\rho_A (\ln |a|^2 \quad 0 \\ 0 \quad \ln |b|^2)) = -\text{Tr}(\ln |a|^2 |a|^2 \quad 0 \\ 0 \quad |b|^2 \ln |b|^2)$
 $= -|a|^2 \ln |a|^2 - |b|^2 \ln |b|^2$

For $a=b = \frac{1}{\sqrt{2}}$ $S_{vN} = +\frac{1}{2} \ln 4 + \frac{1}{2} \ln 4 = 2 \ln 2 \neq 0$! } What is the
 For $a=1, b=0$ $S_{vN} = 0$ } difference?

For $a=b = \frac{1}{\sqrt{2}}$ $|\psi\rangle = \frac{1}{\sqrt{2}} |\uparrow\rangle_A |\downarrow\rangle_B + \frac{1}{\sqrt{2}} |\downarrow\rangle_A |\uparrow\rangle_B$ A is entangled w/ B
 $a=1, b=0$ $|\psi\rangle = |\uparrow\rangle_A |\downarrow\rangle_B$ A is not entangled w/ B

Note that after tracing over B, what we know is $\{p_i, |i\rangle_A\}$ very similar to the statistical case.

A big difference is $S_E(A) + S_E(B) \geq S_E(A \cup B)$, e.g.

